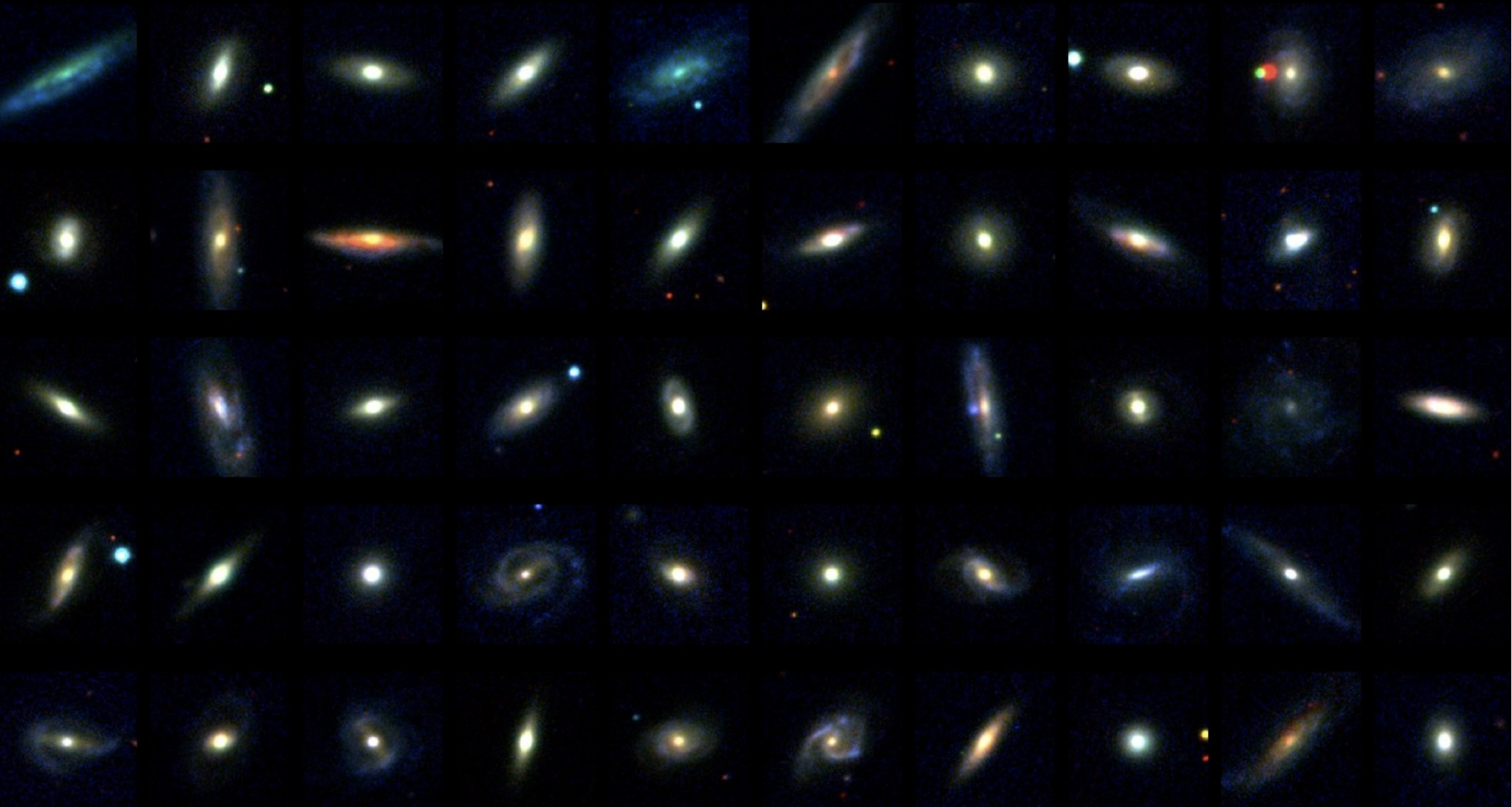




GAMA-SIGMA: Single-Sérsic Galaxy Fitting



Lee Kelvin, ICRAR, March 2011



Survey Regions

GAMA 09

GAMA 12

GAMA 15

Redshifts

SDSS

UKIDSS

Galaxy And Mass Assembly (GAMA) Hubble Tuning-Fork

G00609396

G00508412

G00272822

Groups/
Mergers

G00238211

G00303096

G00560238

G00230636

G00572402

Unbarred Spirals

Ellipticals

G00055143

G00230675

G00047674

GAMA



G00238952

G00600168

Irregulars

Edge-on
Spirals

G00583427

G00484729

G00611220

G00383312

G00240202

Barred Spirals

G00229306

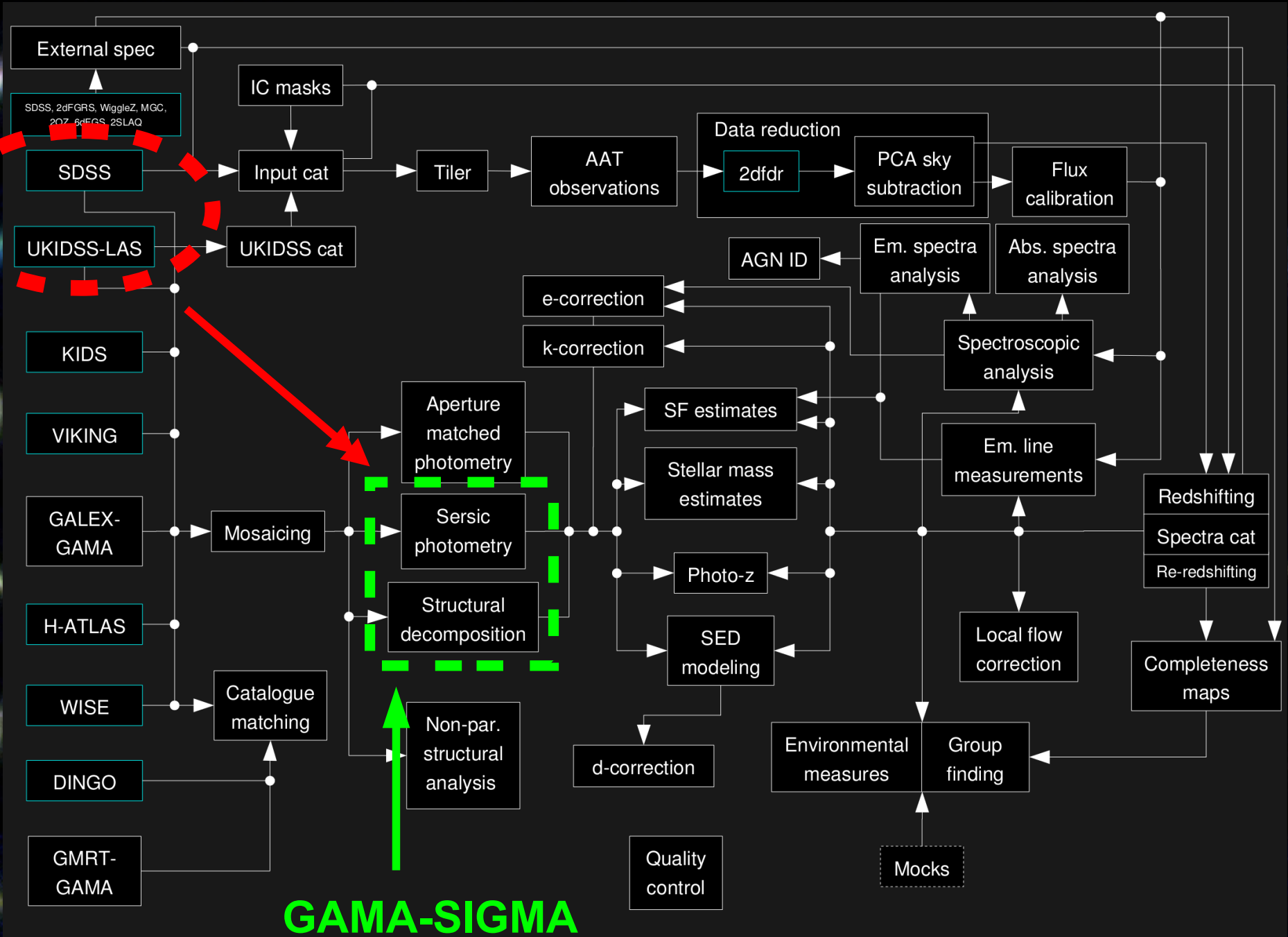
G00619104

G00493832

G00537361

Kru 40"×40"

Dataflow





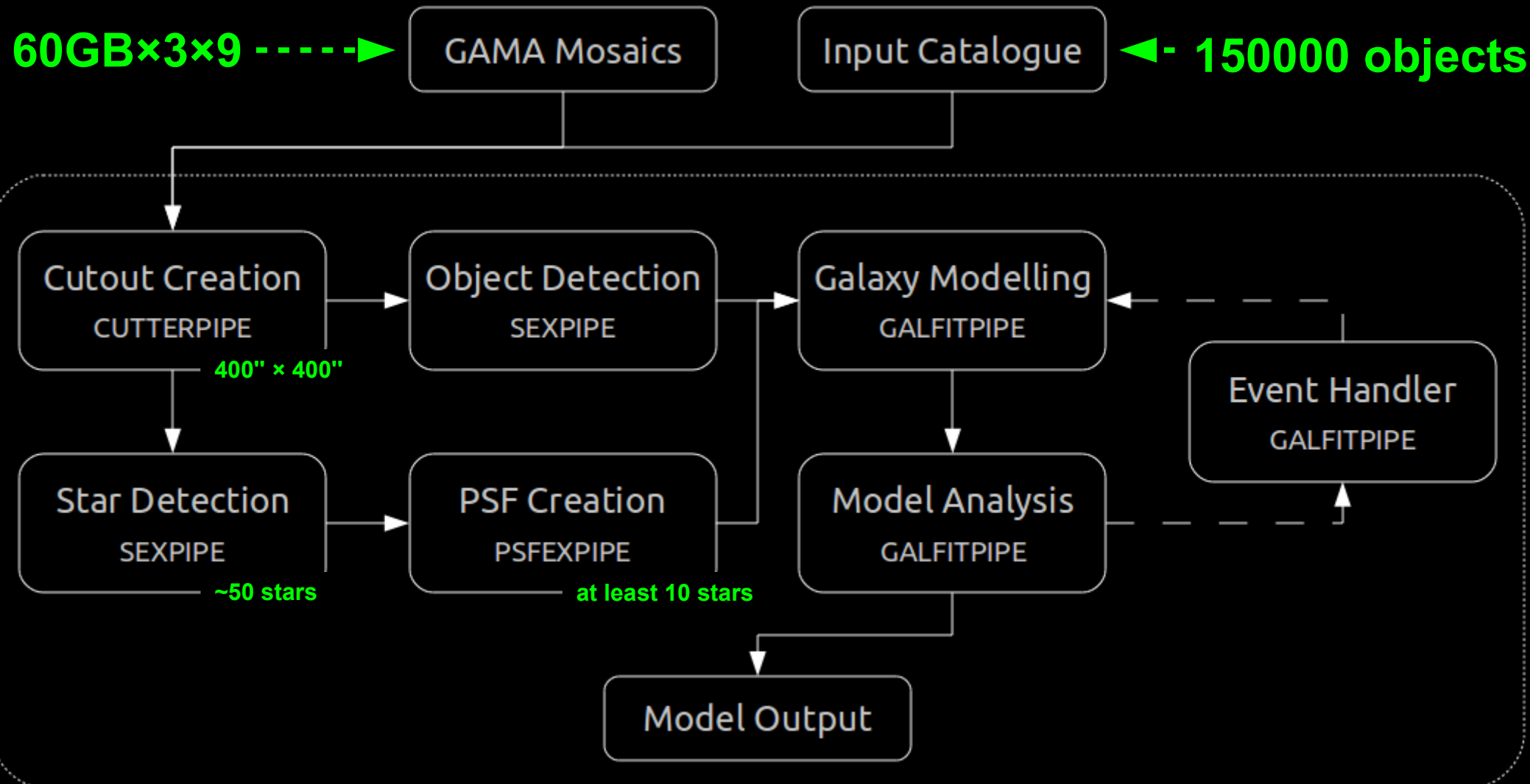
GAMA-SIGMA



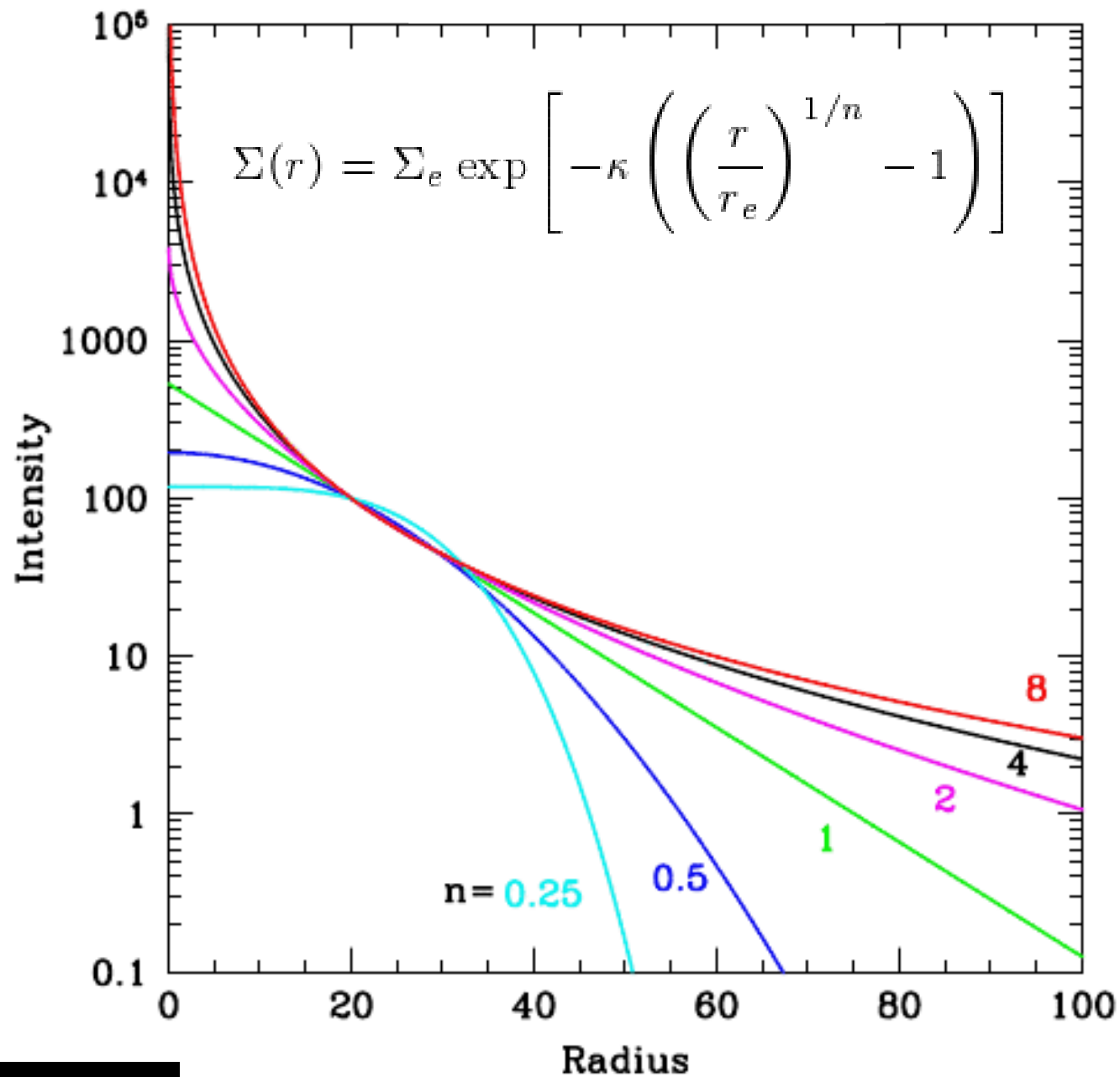
SIGMA: Structural Investigation of Galaxies via Model Analysis

→ Written in R

→ Run time: 20 sec!



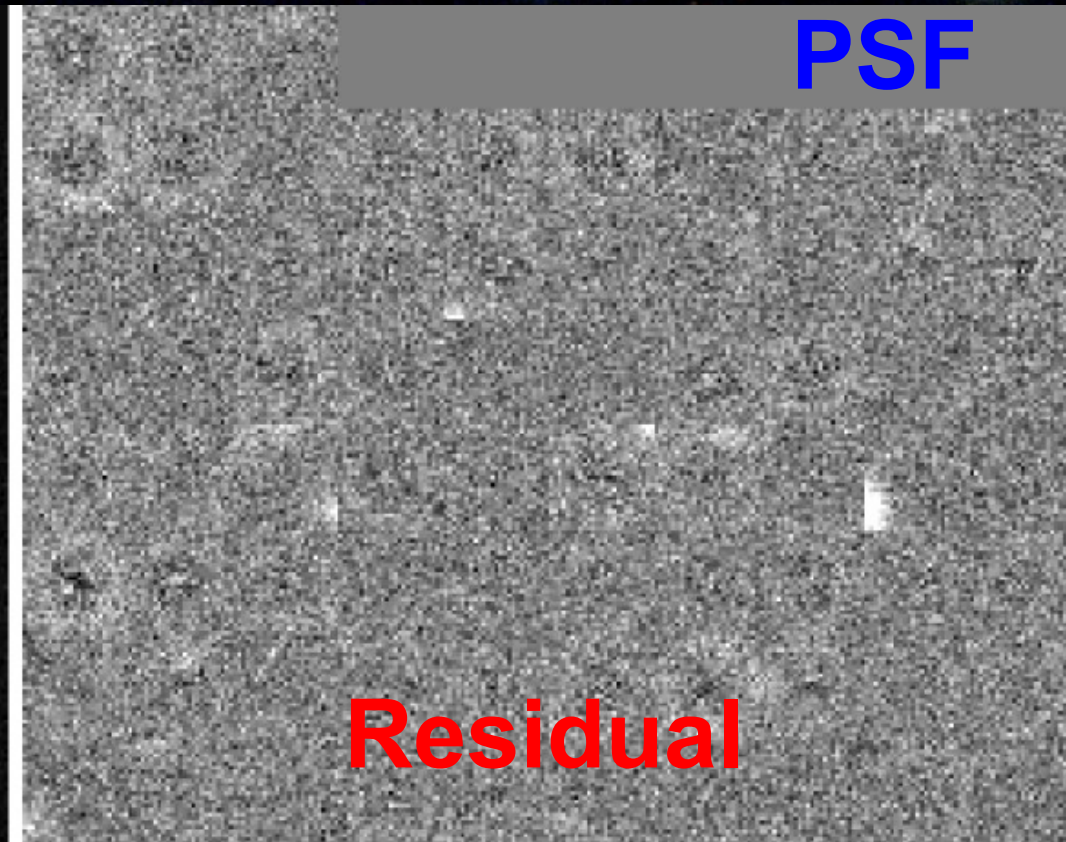
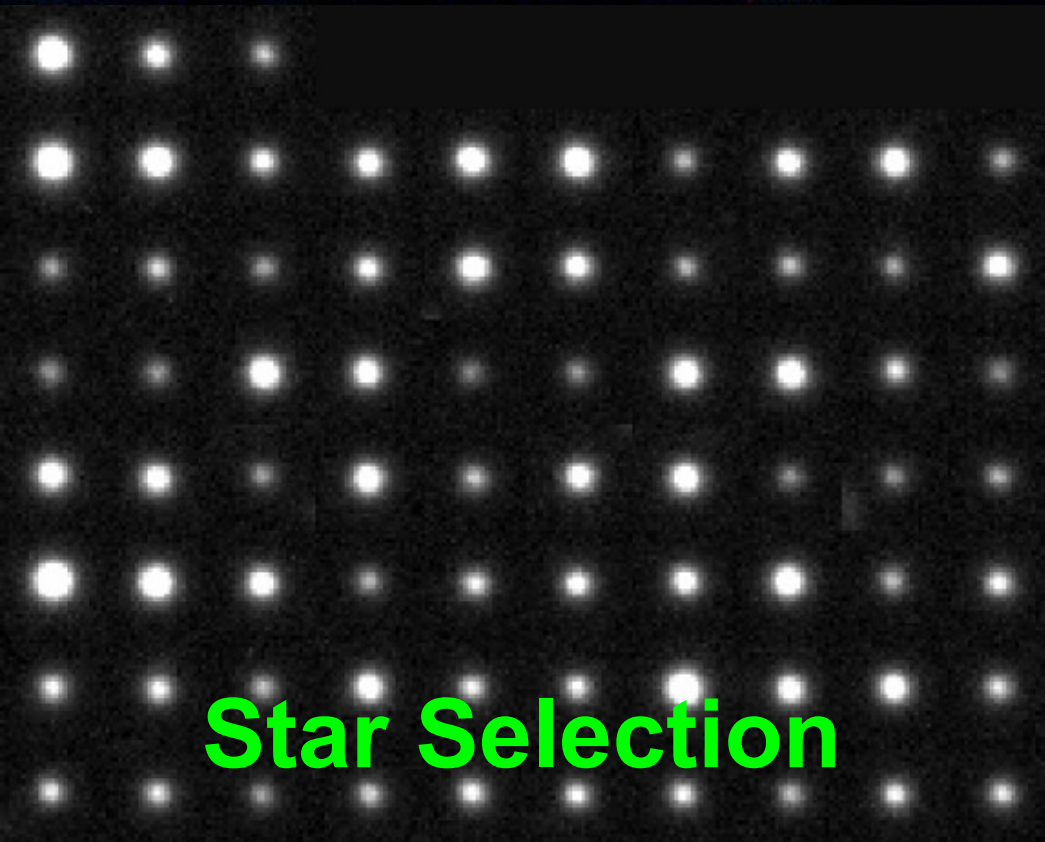
The Sérsic Profile



Jose Luis Sérsic, 1963

PSF Extractor

Empirical point-spread function creator



Pros:

- Extremely Fast
- High fidelity PSF's

Cons:

- Short on documentation

PSFEx: Emmanuel Bertin

GALFIT 3.0

- Galaxy fitting algorithm that fits 2D parameterised, axisymmetric functions to images.

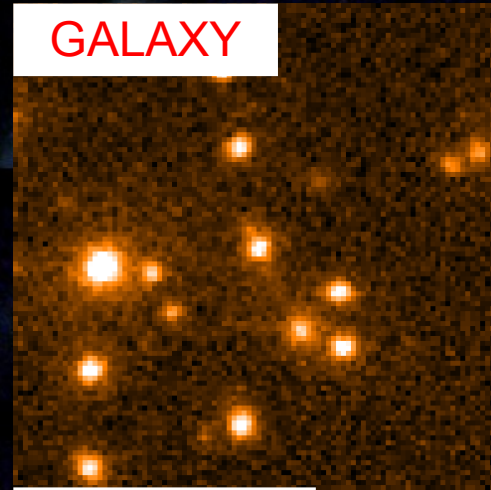
Functions;

- Exponential
- de Vaucouleurs
- Sérsic
- Ferrer / Nuker
- PSF (& Moffat)

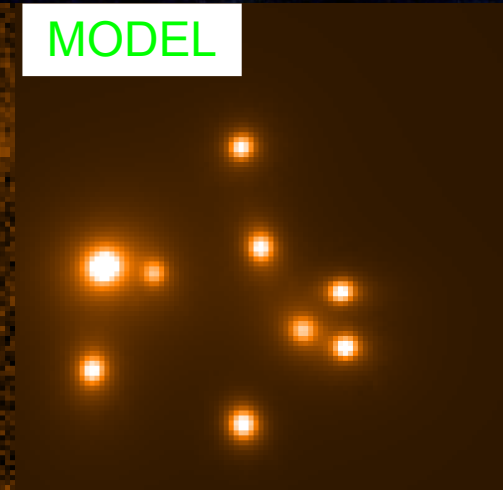
Able to vary;

- Magnitude
- Sérsic Index / Power
- R_e / R_s
- Position angle
- Ellipticity

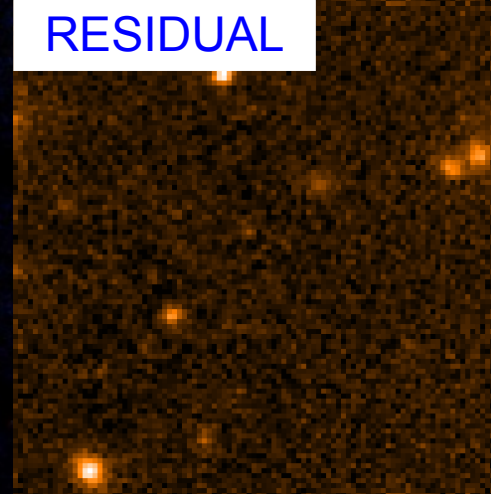
GALAXY



MODEL



RESIDUAL



G00198157_r

RA: 138.46 DEC: -0.76
z: 0.341

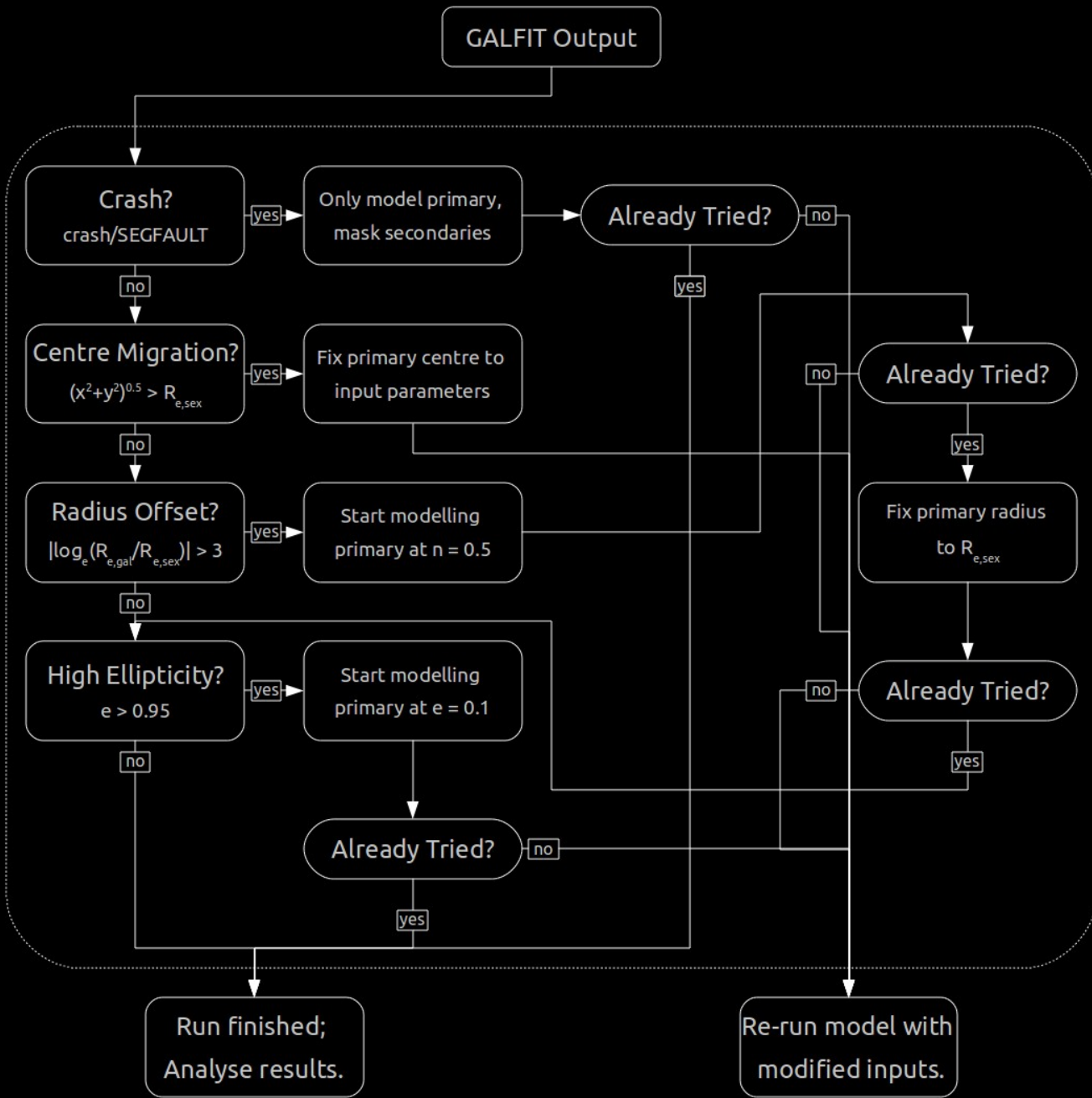
SIGMA Sérsic mag: 19.3
SDSS Petro mag: 19.0

r_g : 1.5" n: 1.8

GALFIT: Chien Peng



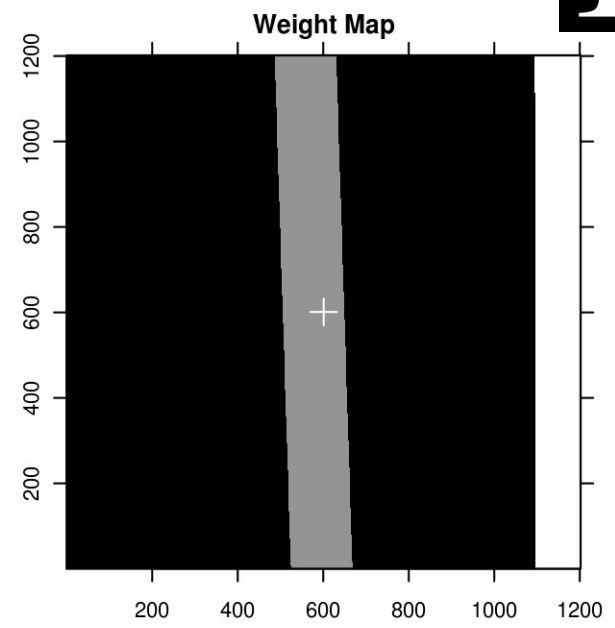
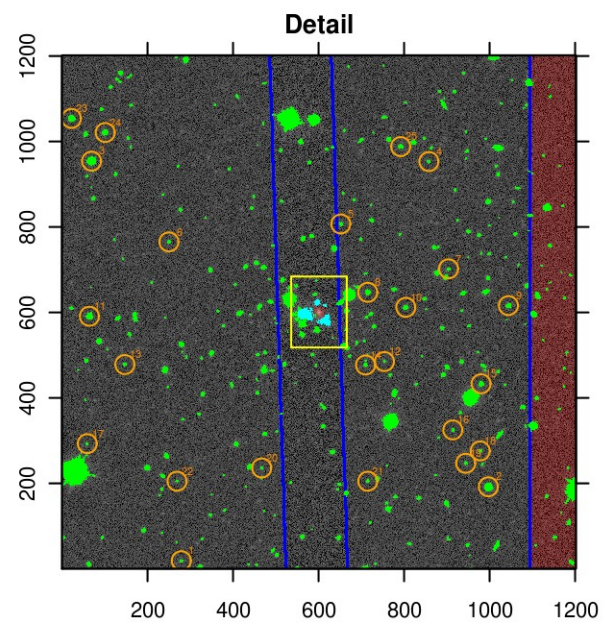
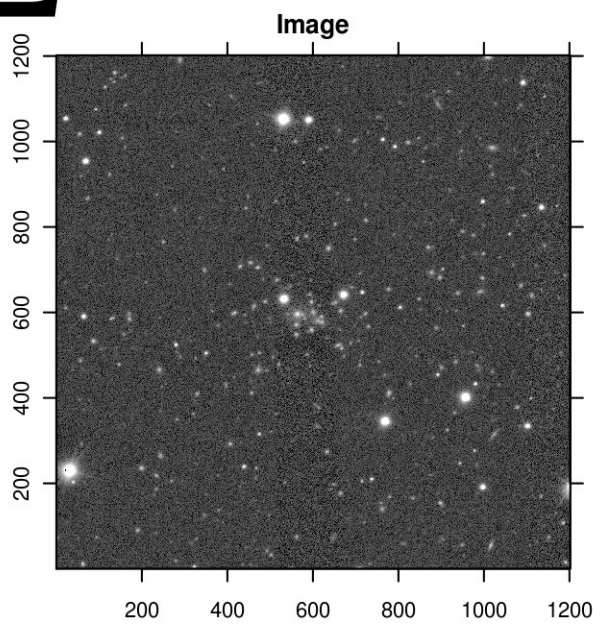
GAMA-SIGMA



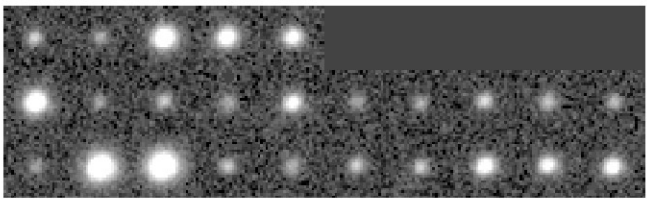
- * Scan for problems
- * Attempt to fix them
- * Assign a binary history flag
- * Assign a quality flag (eg - high sersic index, bad χ^2).



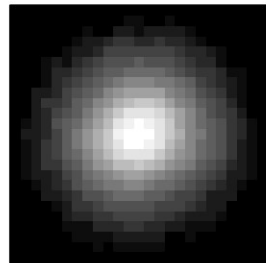
GAMA-SIGMA



PSF Star Sample



PSF



PSF Star Residuals

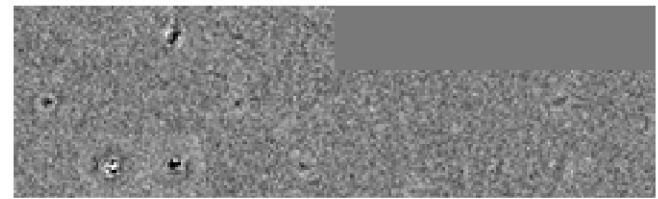
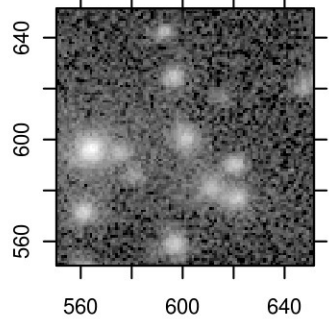
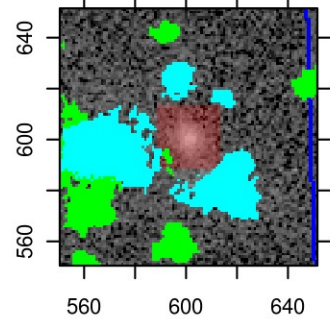


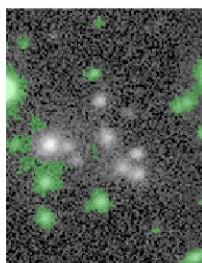
Image (100 pix)



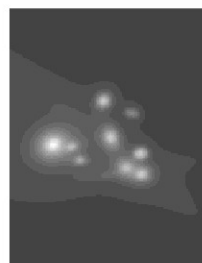
Detail (100 pix)



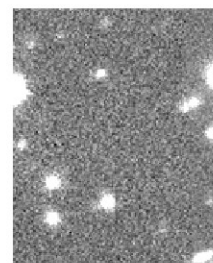
Fitting Region



Model

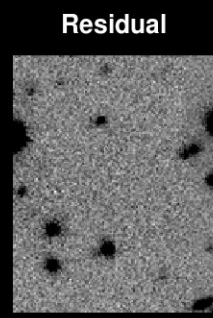
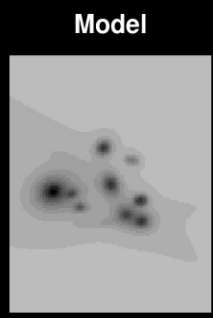
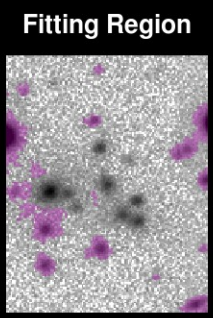
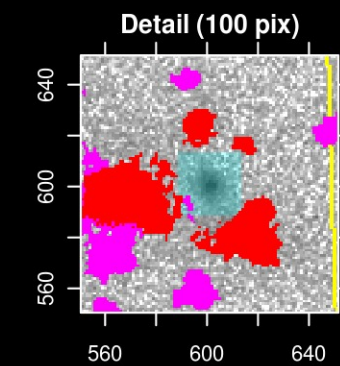
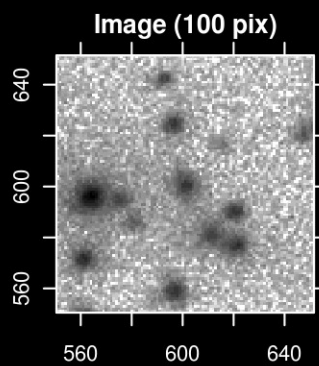
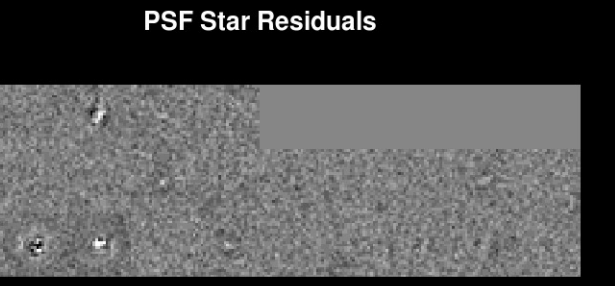
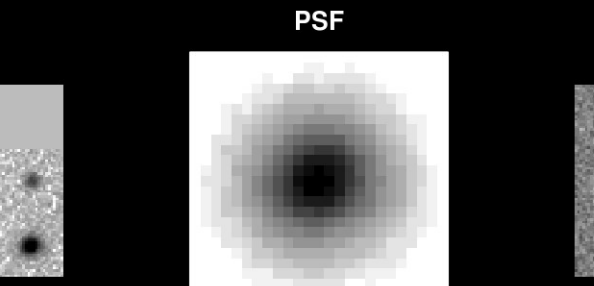
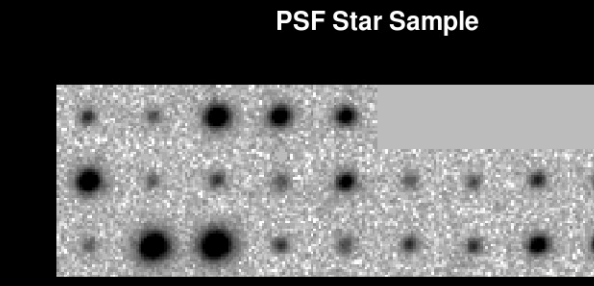
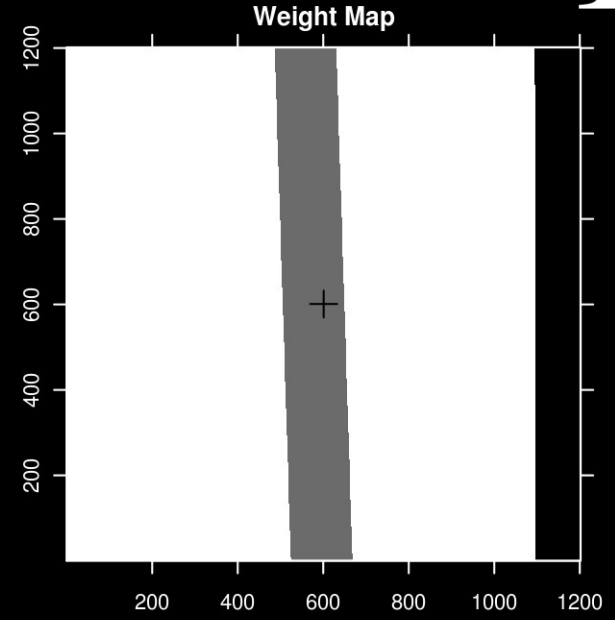
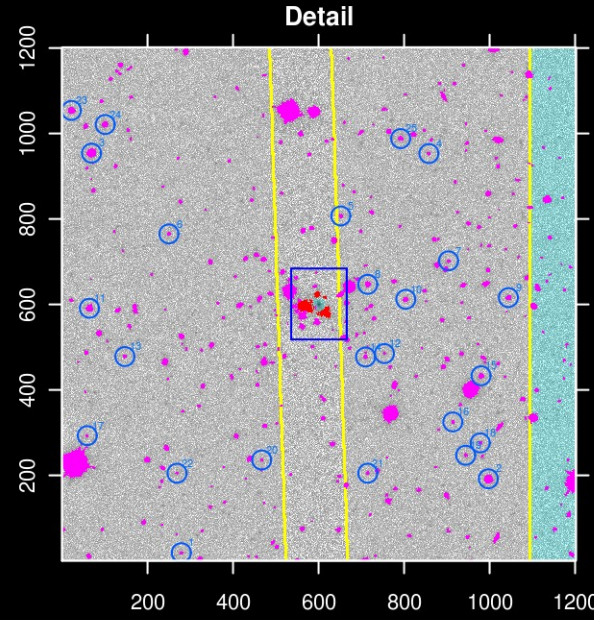
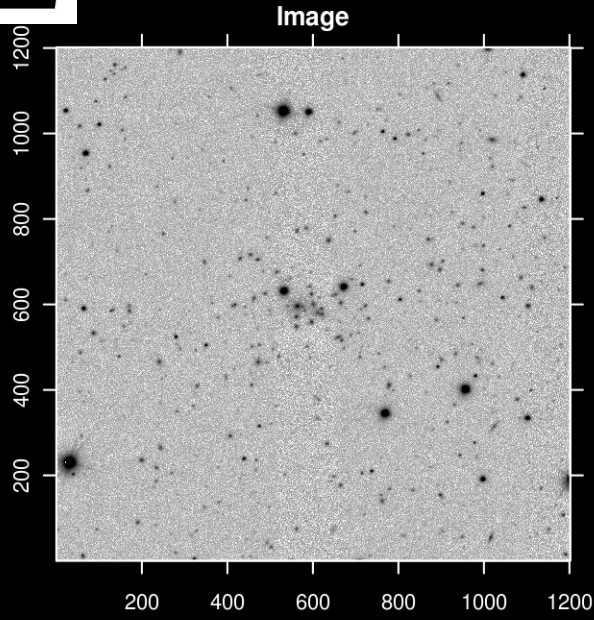


Residual



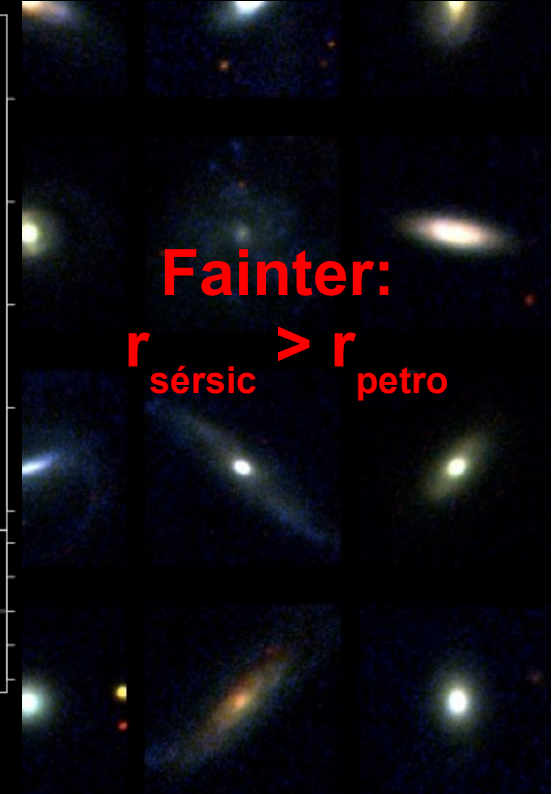
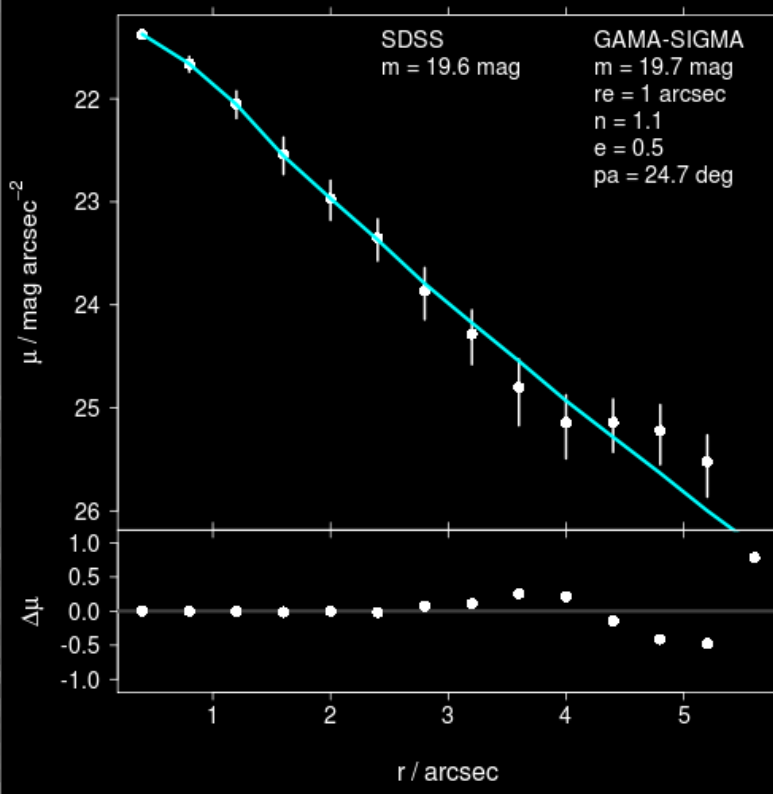
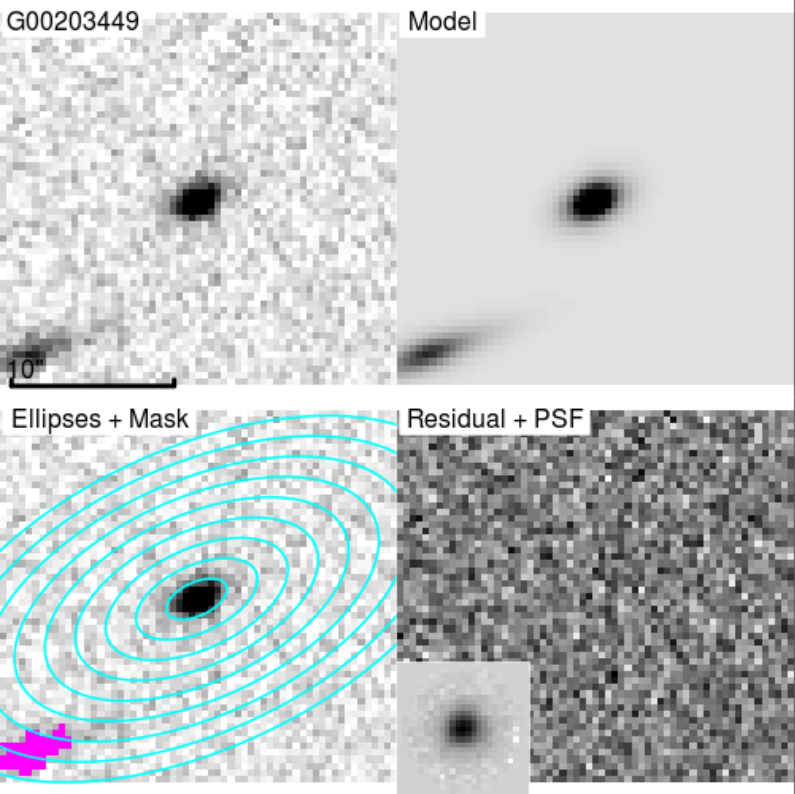
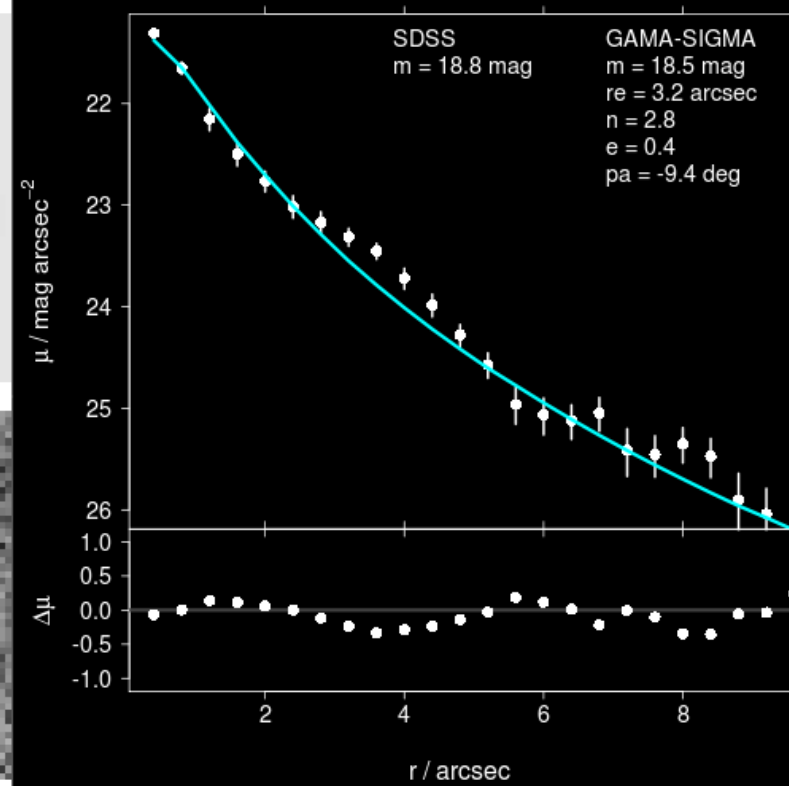
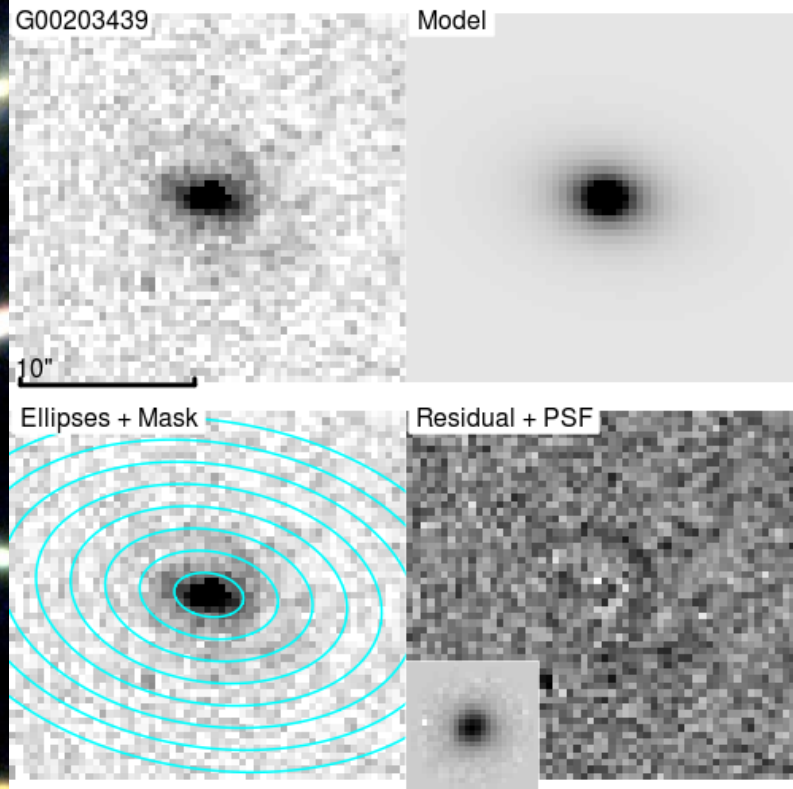


GAMA-SIGMA

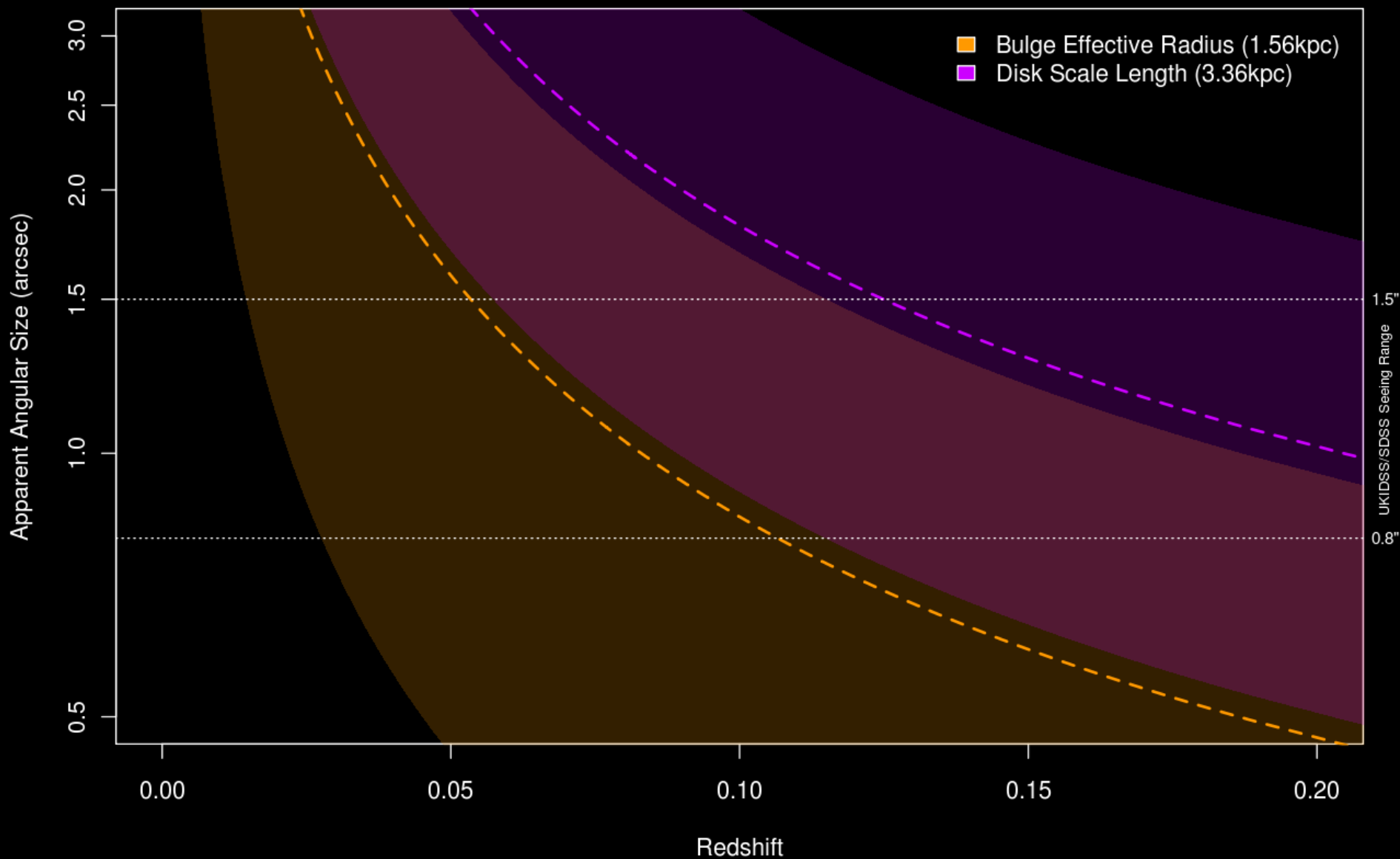


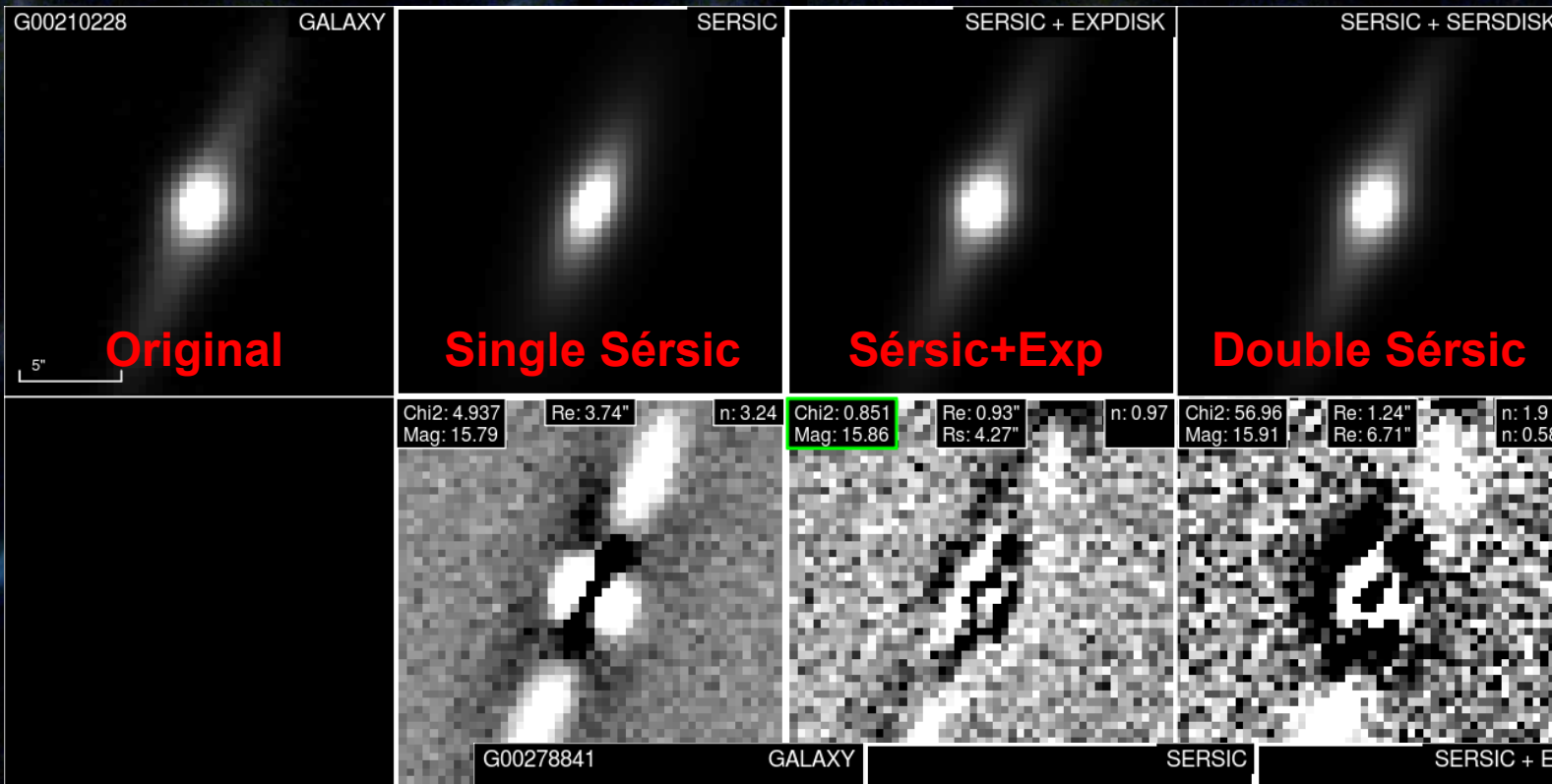
Single-Component

Brighter:
 $r_{\text{sérsic}} < r_{\text{petro}}$

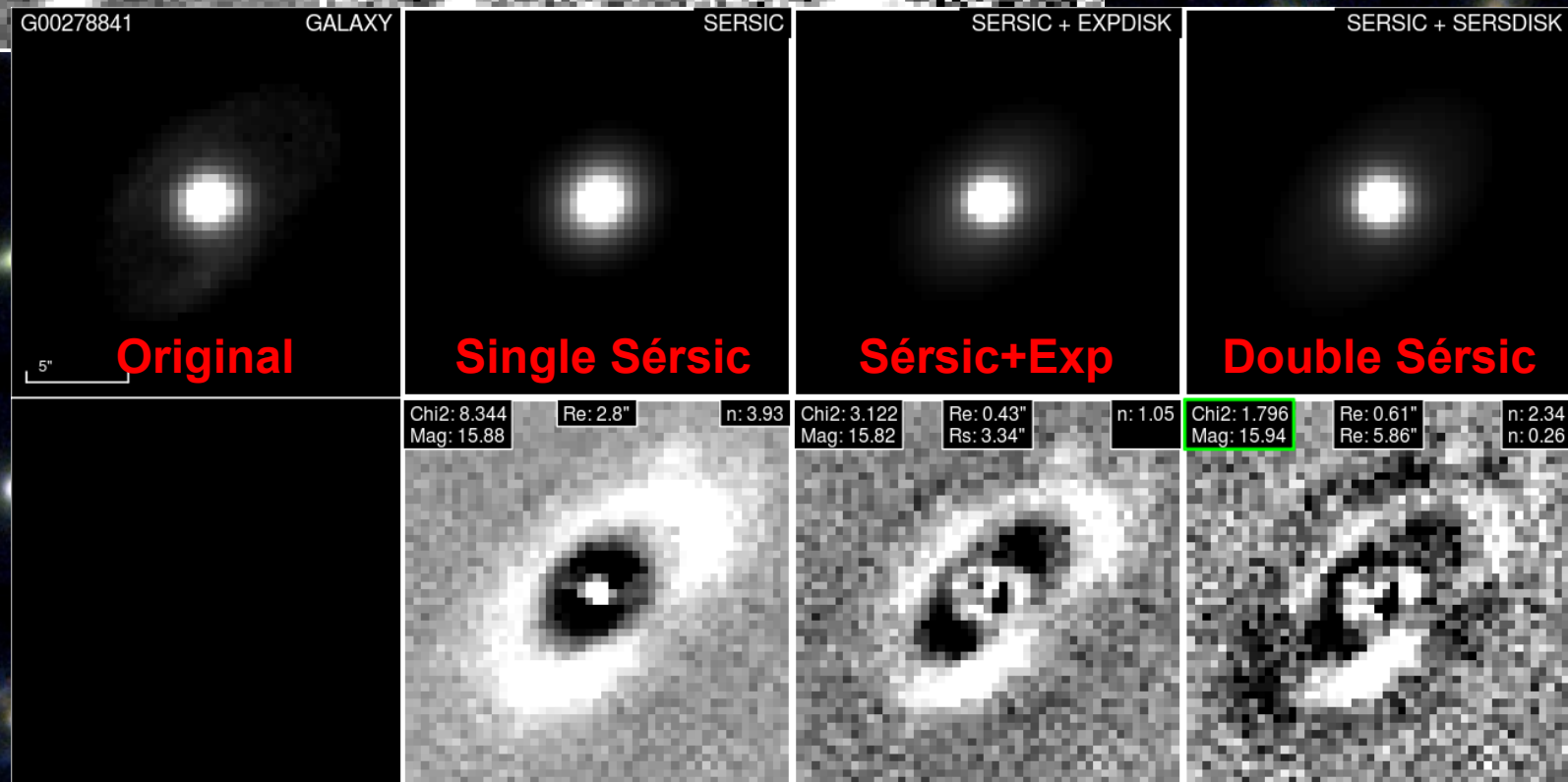


Multi-Component Modelling





Multi-Component



Pros:

- Easy to model multiple components

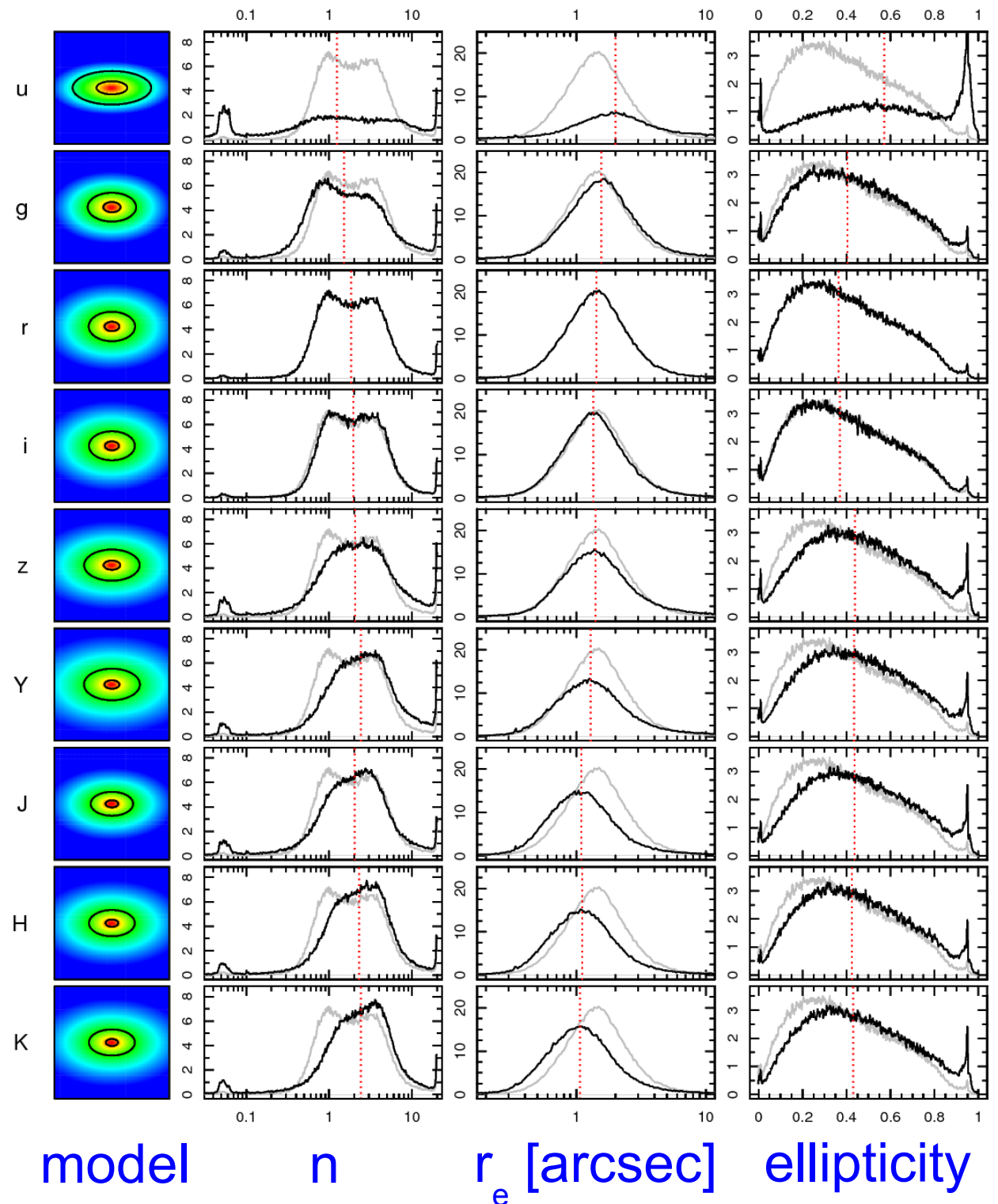
Cons:

- Tricky choosing correct fit in an automated fashion

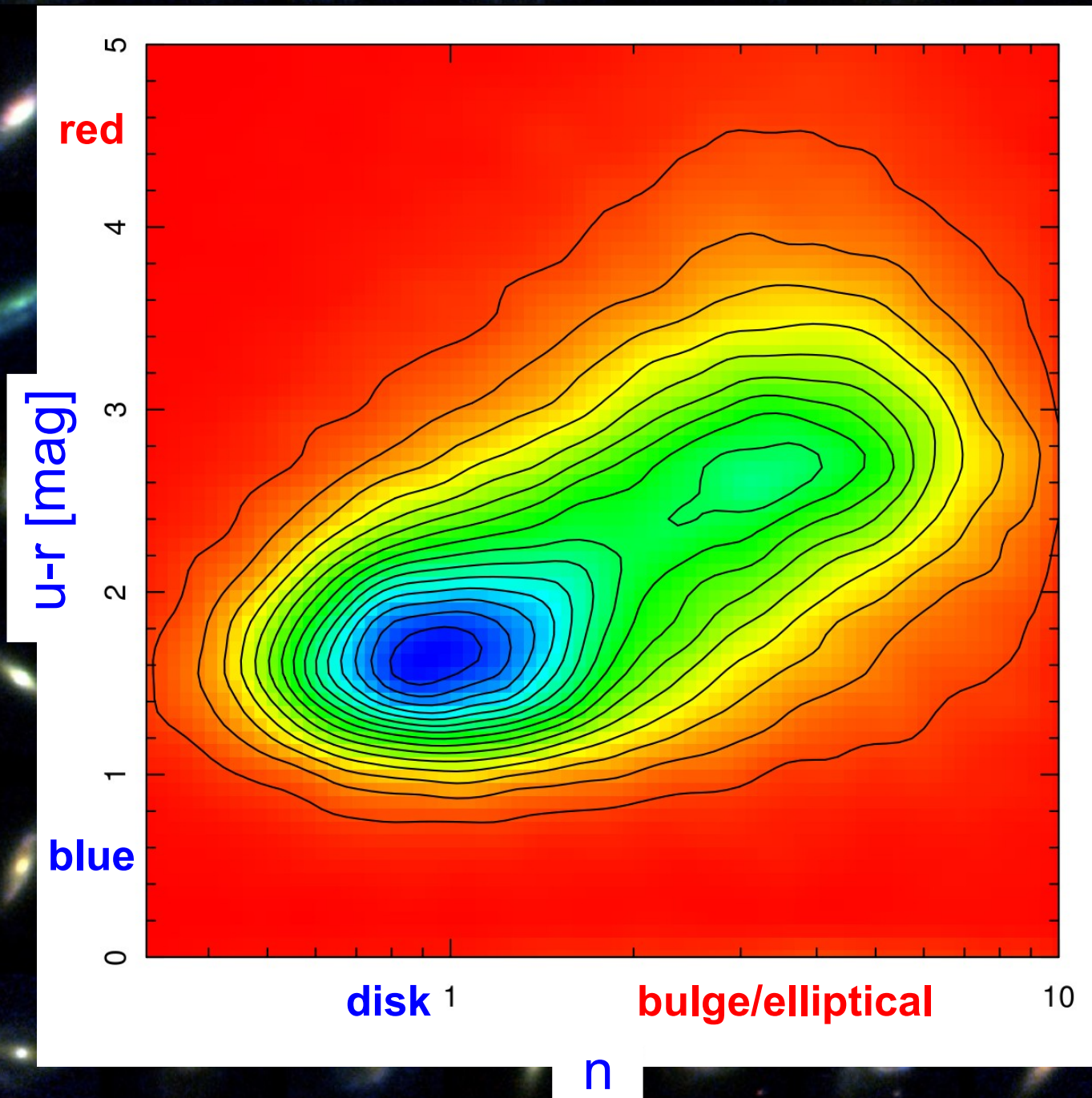
Global Outputs

$u \rightarrow K$:

- increasing n
- decreasing r_e
- stable e ($1-b/a$)



Sérsic Index – colour relation

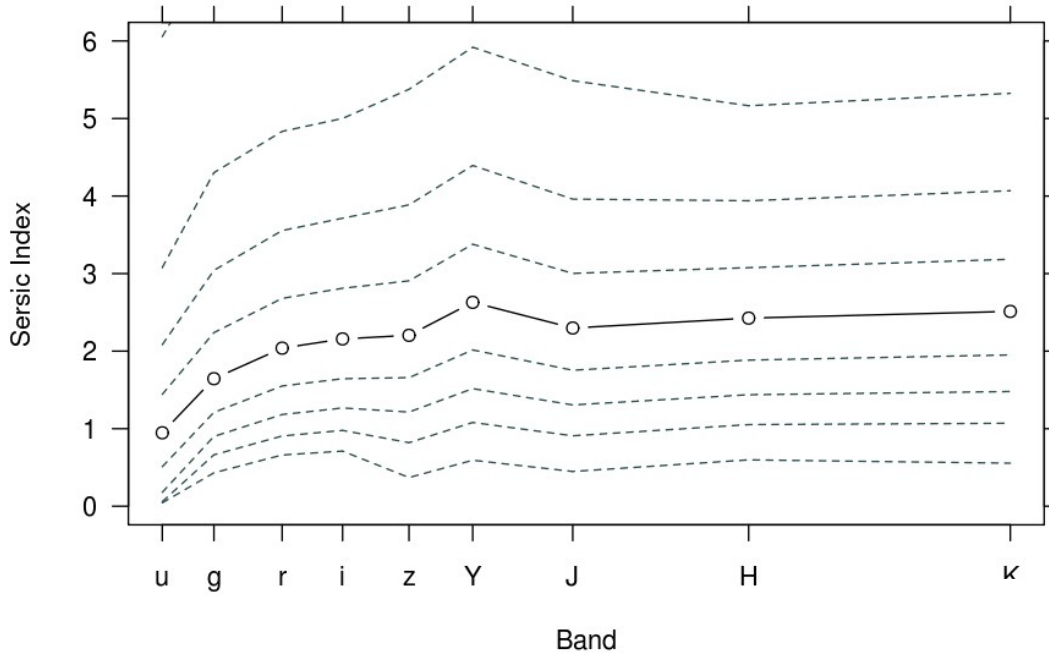


Bimodal:

$> n=1$ (disk-like)

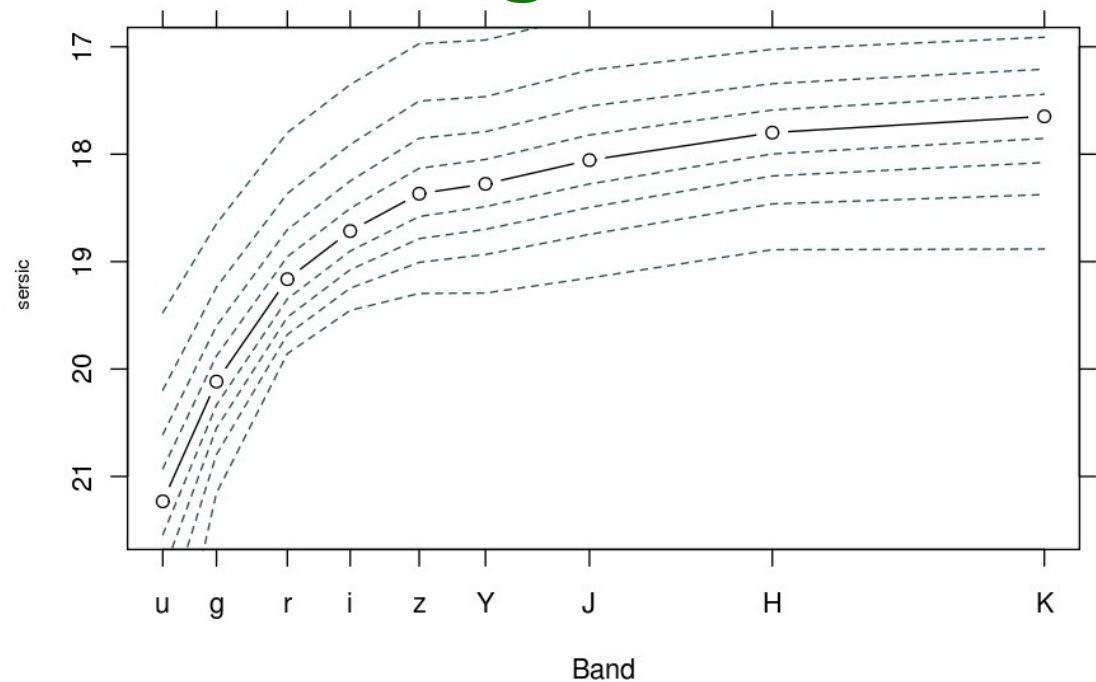
$> n=4$ (spheroid-like).

Sérsic Index



Wavelength dependency on recovered parameters

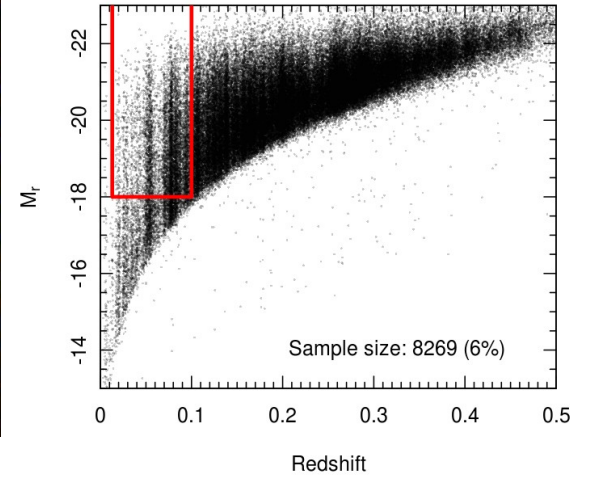
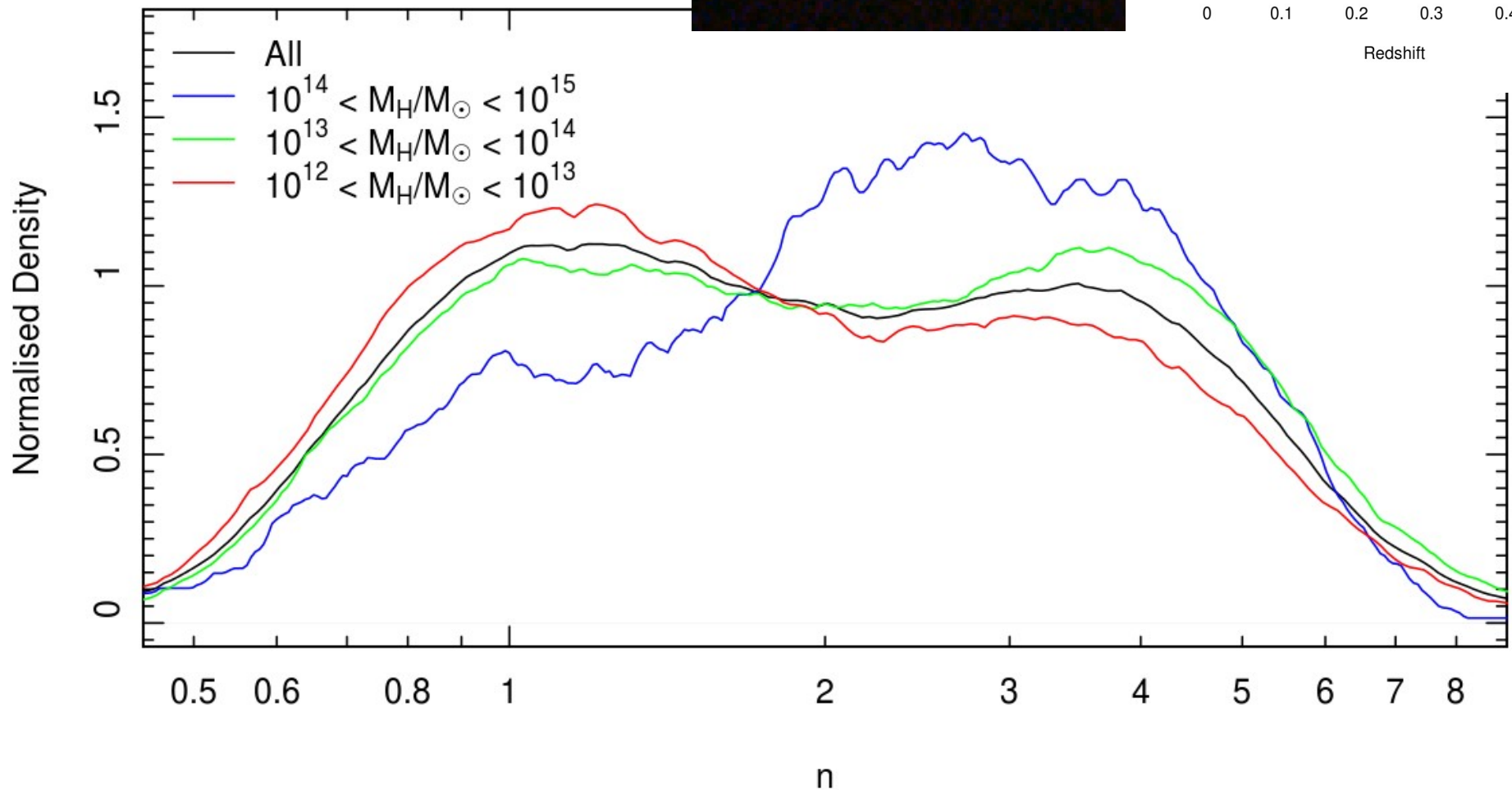
Magnitude



'Predictions' can be calculated to convert into other bands

Cool Plots

Sérsic Index - Halo Mass



Halo Masses: Aaron Robotham

Cool Plots

Halo Masses

· Aaron Robotham

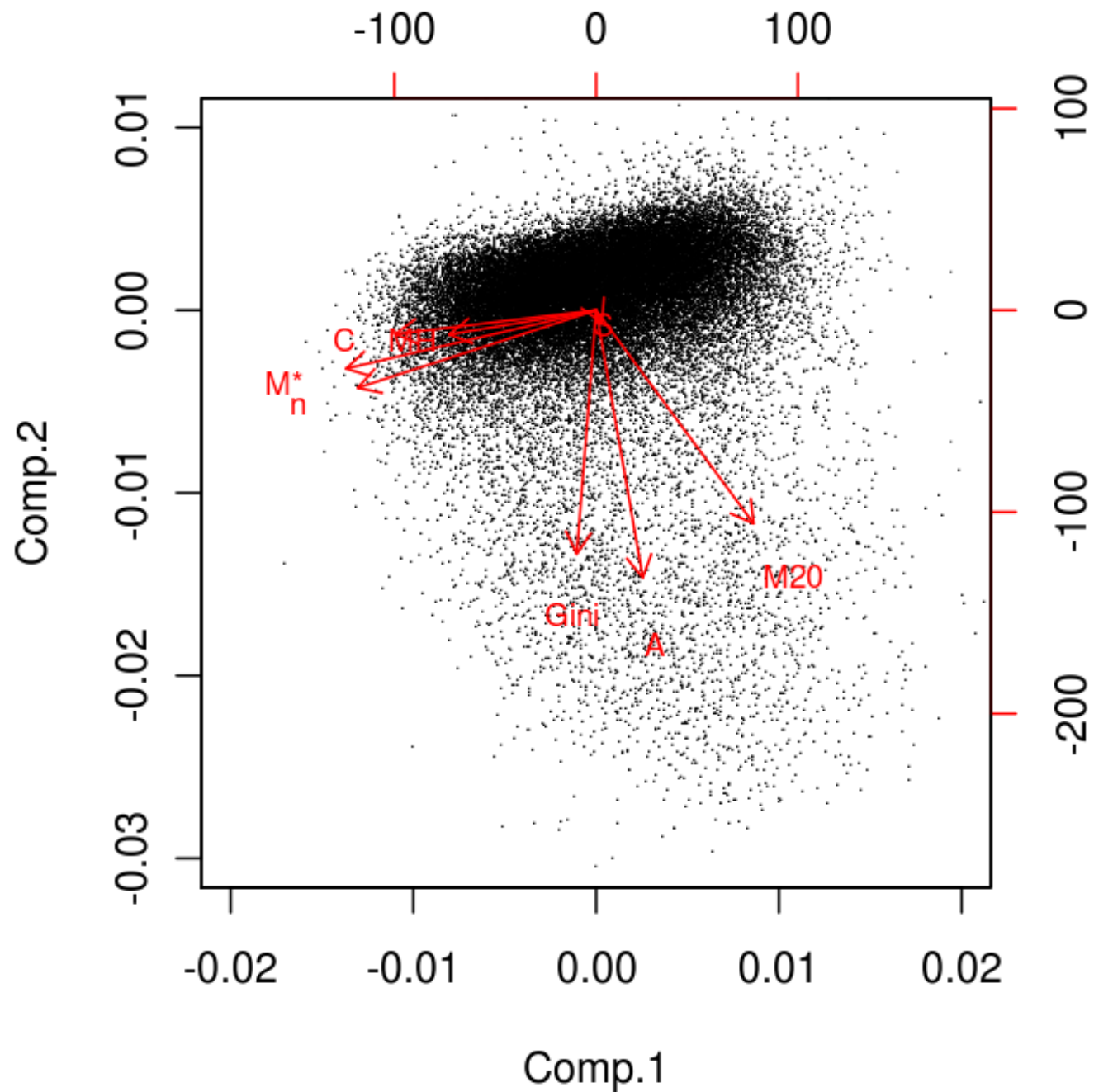
CAS/Gini/ M_{20}

· Chris Conselice

Stellar Masses

· Ned Taylor

Principle Component Analysis



Summary

- Single-Sérsic photometry recovers flux that traditional methods cannot.
- Multi-component provides even better results but there are limitations.
- Important to use multi-wavelength data and find an improved dust model.

What Now?

SersicCat:

- A high-fidelity sample of $\sim 150,000$ galaxies with measurements of effective radii, sérsic indices, sérsic magnitudes and other physical properties.
- $> 95\%$ success rate
- Awaiting QC approval!

What Next?

- Full structural decomposition!
 - Which multi-component model is the 'best'?
- Exploration of how physical properties (e.g.: Sérsic index, size) relate to Stellar Mass/Environment/Morphology.
- Study of the evolution of spheroidal and disk like structures – do bulges have a similar formation mechanism to ellipticals?